Thirty Meter Telescope Adaptive Optics System Error Budgets and Requirements Traceability

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ABSTRACT

The Thirty Meter Telescope (TMT) uses error budgets to understand and track the expected science parameters of the Observatory. In this paper, we demonstrate how the top-level requirements for wavefront error in both Multi-Conjugate Adaptive Optics (MCAO) and Natural Guide Star Adaptive Optics (NGSAO) modes have been decomposed and allocated between various sources that may cause performance degradation. We also show how those values have been integrated into the requirements for each individual subsystem. By integrating these error budgets into our requirements management process, we are able to maintain traceability between science and design, and understand how changes at a low-level could affect the overall AO performance of the Thirty Meter Telescope.

Keywords: TMT, NFIRAOS, WFE, Budget, Requirements, Traceability

1. INTRODUCTION

The TMT NFIRAOS NGSAO and LGS MCAO Wavefront Error Budget¹ takes the science requirements (SRD) for wavefront error and breaks them down following the TMT system decomposition. Observatory architecture requirements (OAD) are managed by Systems Engineering, and they are further decomposed into lower-level requirements which are managed by each subsystem team. The budget is split into LGS MCAO mode (Table 1) and NGSAO mode (Table 2).

To reduce the number of requirements and facilitate future changes, a single requirement is used for subsystems that have the same WFE allocation regardless of whether LGS MCAO (on-axis) or NGSAO mode. These allocations are shown in gray in the NGSAO budget, which, if also gray in the LGS on-axis budget, also denotes commonality between the LGS MCAO on-axis, 17"x17", and 30" field of view allocations. Some requirements in Table 1 are shown in pink, denoting common terms that, if changed, would require an update to the telescope seeing-limited budget. Values shown in white are for any allocations specific to LGS MCAO mode.

1.1 LGS MCAO Budget

The LGS MCAO budget (Table 1) is split into High Order and Low Order modes, with each showing the requirements decomposition based on on-axis, 17"x17", or 30" (nm) field of views. The higher order modes contain all modes beyond tip, tilt, plate scale, and global focus. Low order modes include global tip/tilt, focus, and plate scale that are controlled only by low order NGS WFS.

1.2 NGSAO Budget

The NGSAO budget (Table 2) is split into High Order and Low Order modes, and the requirements are decomposed based on mR=8 or mR=12 guide stars. The low order modes include tip/tilt only.

		On axis	17"x17"	30" Diameter
REQ #	Terms NFIRAOS LGS MCAO and IRIS WFE	L0 L1 L1 L2 L3 187	L0 L1 L1 L2 L3 191	L0 L1 L1 L2 L3 203
REQ-0-SRD-0820				
REQ-1-ORD-3530 REQ-1-ORD-3532	High Order Modes	173	190	190
	Telescope			
REQ-1-OAD-0251	TCS Rupil misragistration (Control)	6	6	6
REQ-1-OAD-0252	M1S	29	29	29
	M1 static shape	29	29	29
REQ-1-OAD-0253	M1CS Segment dynamic misalignment	14	14	14
REQ-1-OAD-0254	M2S	13	14	16
	M2 Static Shape	11	11	11
	Pupil misregistration (M2 actuators)	6	6	6
REQ-1-OAD-0255	M3S	11	11	11
	M3 Static Shape Pupil microgistration (M3 actuators)	9	9	9
REQ-1-OAD-0256	APS	16	16	16
	M1 shape calibration	16	16	16
REQ-1-OAD-0257	Facilities	30	30	30
	Dome Seeing	22	22	22
	Mirror Seeing	20	20	20
REQ-1-OAD-0258	Instrumentation NFIRAOS SYSTEM	157	176	176
REQ-1-OAD-0259	NFIRAOS OM	50	58	60
	NFIRAOS-to-Telescope misalignment	0	20	20
	Uncorrectable error NCPA calibration error	35	35	35
	DM/WFS pupil distortion	12	12	12
	DM/WFS pupil misregistration	16	16	16
	Telescope pupil misregistration (Measurement error)	6	6	6
	Dynamic higher order error	5	5	5
	Output beam misalign	15	15	15
REQ-1-OAD-0260	AO Comp: WC	51	51	51
	Failed actuators	19	19	19
	Hysteresis	20	20	20
	Dynamics	11	11	11
REQ-1-OAD-0261	Surface flattening	42	42	42
REQ-1-OAD-0261	AO Comp: LGSWFS	44	44	44
	Offset/gain calibration	14	14	14
	Pt. src tomographic approx	0	0	0
	Rayleigh fratricide	4	4	4
	Signal variability	23	23	23
	Diff. atmospheric refractive index Chromatic anisoplanatism	1/	1/	0
	Lenslet throughput and aberrations	28	28	28
REQ-1-OAD-0262	AO Comp: RTC	28	28	28
	Numerical precision Cn2 Profile	20	20	20
REQ-1-OAD-0263	AO Architecture	130	148	148
	DM fitting error	74	74	74
	DM projection error	48	85	26
	Tomography Error	48	53	53
	TMT pupil Function	14	4	4
	LGS WES non-linearity	18	1/	17
	LGS WFS noise	51	53	53
	Simulation Undersampling	48	48	48
REQ-1-0AD-0264	IRIS Design residuals	40	40	40
	Chromatic aberration	14	14	14
	Fabrication/installation	10	10	10
	Alignment accuracy Cooldown	8	8	8
	Surface quality	26	26	26
	Dynamic higher-order error	3	3	3
	ADC effects Glass inhomogeneities	4	4	4
	NCPA calibration error	10	10	10
	Others	14	14	14
REQ-1-0AD-0265	LGSF Surface roughness	34 30	34	34
	Alignment and Fabrication	15	15	15
REQ-0-SRD-0850 REQ-1-ORD-2730	Low order Modes (Tip/tilt, Focus and Plate Scale)	68	68	68
neq I ono Erso	Telescope			
REQ-1-OAD-0266	STR, M1, M2 and M3	37	37	37
	Windshake tip/tilt error Windshake plate scale error	16	16	16
	Telescope structure vibration	28	28	28
1	Telescope tracking jitter	17	17	17
RE0-1-04D-0267	Instrumentation	54	E4	EA
REQ-1-OAD-0267	NFIRAOS OM	22	22	22
REQ-1-OAD-0269	Internal NFIRAOS vibration	10	10	10
	Field dependent WFE	20	20	20
	AU Comp: WC TTS/DM dynamics	0	0	0
	DM hysteresis	0	0	0
REQ-1-OAD-0270	AO Comp: RTC/RPG	0	0	0
BEO-1-04D-0271	RTC/RPG implementation	0	50	50
ncc-1-0AD-02/1	Turbulence tip/tilt	32	32	32
	Turbulence plate scale	38	38	38
REQ-1-OAD-0272	IRIS NEIRAOS to IRIS vibration	16	16	16
	OIWFS to Imager vibration	10	10	10
	Internal IRIS imager vibration	7	7	7
	WFS (OIWFS/ODGW)	0	0	0

Table 1. TMT NFIRAOS LGS MCAO Wavefront Error Budget.

REO #	Tarme	10	_	mR=	8 guide	star	12	10	mR=1	L2 guide	star	12
REQ-0-SRD-880	Terms	10				12	1.5	105		u	LZ	13
REQ-0-SRD-881	NFIRAOS NGSAO and IRIS WFE	158	_	_				185	_			
REQ-1-ORD-3671	High Order Modes		1	49					185			
	Telescope			_					_			
REQ-1-OAD-0251	TCS Pupil misregistration (Control)			-	6	6				6	6	
REQ-1-OAD-0252	M1S		_		29					29	-	
BEO 1 040 0353	M1 static shape			_		29					29	
REQ-1-0AD-0253	Segment dynamic misalignment			_	14	14			_	14	14	
REQ-1-OAD-0254	M2S				13					13		
	M2 Static Shape					11					11	
	Pupil misregistration (M2 actuators)					6			_		6	
REQ-1-OAD-0255	M3S				11					11		
	M3 Static Shape Pupil misregistration (M3 actuators)		_			9					9	
REQ-1-OAD-0256	APS			_	16	0			_	16	0	
	M1 shape calibration					16					16	
BEO-1-04D-0257	Facilities			-	30					30		
heq I one dest	Dome Seeing			_	50	22			_	50	22	
	Mirror Seeing					20					20	
BEO-1-0AD-0273	Instrumentation NEIRAOS SYSTEM				134					174		
REQ-1-OAD-0274	NFIRAOS OM		_			51					51	
	NFIRAOS-to-Telescope misalignment						0					0
	Uncorrectable error						35					35
	Registration Drifts after Calibration						15					15
	Image Quality at Pyramid tip						25					25
	(Measurement error)						6					6
	Dynamic higher order error						5					5
	Output beam misalign					54	15		_		54	15
REQ-1-0AD-0260	AC Comp: WC Actuator saturation					51	0				51	0
	Failed actuators		_				19					19
	Hysteresis						20					20
	Influence function						0					0
	Surface flattening		_				42					42
REQ-1-OAD-0275	AO Comp: PWFS WFS					38	45		_		38	45
	Pupil image location						15					15
	Imperfect pyramid						16					16
	Pupil image quality						16					16
	Pupil image distortion						16					16
	Modulation errors					_	0					0
REQ-1-OAD-0276	AO Comp: RTC					20	20				20	20
REQ-1-0AD-0277	AO Architecture					105	20		_		152	20
	DM fitting error						74					74
	PWFS aliasing error						16					16
	Servo Lag						18					18
	WFS non-linearity						64					64
	WFS noise						0					110
REQ-1-OAD-0264	IRIS			_	40		20			40		20
	Design residuals					7					7	
	Chromatic aberration					14					14	
	Alignment accuracy		_			8					8	
	Cooldown					6					6	
	Surface quality Dynamic higher-order error		-			26		-		-	26	
	ADC effects					4					4	
	Glass inhomogeneities					12					12	
	NCPA calibration error					10					10	
REQ-1-ORD-3669	Low Order Modes (Tip/Tilt and Focus)		2	29		14			29		14	
	Telescope		_	_					_			
REQ-1-OAD-0278	STR, M1, M2 and M3 Windshake tin/tilt error		_	_	22	2			_	22	2	
	Telescope structure vibration					21					21	
	Telescope tracking jitter					5					5	
BEO-1-04D-0279	Instrumentation NEIBAOS System				10				_	10		
REQ-1-OAD-0280	NFIRAOS OM				10	10					10	
	Internal NFIRAOS vibration						10					10
REQ-1-OAD-0281	AO Comp: WC TTS/DM dynamics					0	0	-	_	-	0	0
	DM hysteresis						0					0
REQ-1-OAD-0282	AO Comp: RTC/RPG		-	_		0					0	
REO-1-04D-0283	AO Architecture					2	0		_		2	0
120 2 000-0205	Turbulence tip/tilt						2					2
REQ-1-OAD-0272	IRIS		_		16					16	4.5	
	OIWES to Imager vibration			-		10				-	10	
	Internal IRIS imager vibration					7					7	
	WFS (OIWFS/ODGW)					0					0	

Table 2. TMT NFIRAOS NGSAO Wavefront Error Budget.

2. TRACEABILITY TOOLS

TMT Error Budgets are generated in Excel and then integrated into the Rational Dynamic Object Oriented Requirements System (DOORS)³. DOORS is the source for all TMT requirements as the database allows us to link parent/child requirements together while also being able to generate user-friendly requirement documents for each subsystem.

Each top-level wavefront error allocation from the budget is translated into an individual requirement in DOORS. These allocations then flow down to a lower-level requirement for each contributing subsystem. Links are created between the requirements, providing us with full traceability from top to bottom. Traceability reports are generated directly from DOORS or using the DOORS Trace Tree⁴ tool which shows us the same information in a graphical output (Figure 1, Figure 2). The graphical outputs are extremely useful when trying to identify if a requirement has not yet been traced, or has been mis-traced, to parent or child requirements.



Figure 1. NGSAO & LGS MCAO WFE Requirement Traceability for IRIS (note: not all IRIS requirements shown).



Figure 2. LGS MCAO Requirement Flowdown Example from Trace Tree.

3. NEXT STEPS

Work on the TMT NFIRAOS NGSAO and LGS MCAO Wavefront Error Budget is still in progress. Future work will include activities such as evolving the IRIS imager 34"x34" field of view allocations and incorporating point source sensitivity.

4. SUMMARY

Error budgets are an essential Systems Engineering tool to estimate the TMT AO system's future performance. Systems Engineering, with input from subsystem teams, is responsible for tracking changes to the overall AO error budget. By using the error budgets in conjunction with our traceability tools, we are able to quickly assess how changes in one part of the system may impact higher-level requirements, helping us to minimize risks associated with meeting system and science requirements.

REFERENCES

- [1] TMT Observatory NFIRAOS LGS MCAO, NGSAO and IRIS Imager Wavefront Error Budget and Current Best Estimate, TMT.AOS.COR.16.062
- [2] TMT Observatory NFIRAOS LGS MCAO, NGSAO and IRIS Imager Wavefront Error Budget and Current Best Estimate Description, TMT.AOS.TEC.08.015
- [3] IBM Dynamic Object Oriented Requirements System, http://www03.ibm.com/software/products/en/ratidoor
- [4] DOORS Trace Tree, https://trace-tree.tmt.org/